

Problem 10.1

①

Generalized Bose gas

$$a) \quad \langle N \rangle = (2S+1) \sum_{\vec{k}} \frac{1}{e^{\beta(\epsilon_{\vec{k}} - \mu)} - 1} \quad \epsilon_{\vec{k}} = A |\hbar \vec{k}|^z$$

$$= (2S+1) \left(\frac{L}{2\pi}\right)^d S_{d-1} \int dk k^{d-1} \frac{1}{e^{\beta(\epsilon_{\vec{k}} - \mu)} - 1} \quad S_{d-1} \hat{=} \text{surface of } d\text{-dimensional unit sphere}$$

$$k = \frac{1}{\hbar} \left(\frac{\epsilon}{A}\right)^{\frac{1}{z}} \quad dk = \frac{\epsilon^{\frac{1}{z}-1}}{\hbar z A^{1/z}} d\epsilon$$

$$\langle N \rangle = (2S+1) \left(\frac{L}{2\pi}\right)^d S_{d-1} \frac{1}{z} \left(\frac{1}{\hbar A^{1/z}}\right)^d \int d\epsilon \epsilon^{d/z-1} \frac{1}{e^{\beta(\epsilon - \mu)} - 1}$$

$$g(\epsilon) = (2S+1) \frac{L^d}{(2\pi)^d} S_{d-1} \frac{1}{z \hbar^d A^{d/z}} \epsilon^{d/z-1} = VB \epsilon^{d/z-1}$$

b) Set $\mu = 0$

$$\langle N_{\max} \rangle = VB \int_0^{\infty} d\epsilon \epsilon^{d/z-1} \frac{1}{e^{\beta\epsilon} - 1} \quad \begin{array}{l} x = \beta\epsilon \\ d\epsilon = k_B T dx \end{array}$$

$$= VB (k_B T)^{d/z} \underbrace{\int_0^{\infty} dx \frac{x^{d/z-1}}{e^x - 1}}_I$$

BE condensation if integral converges

upper bound: integrand $\sim e^{-x} \Rightarrow$ converges

lower bound: $\int_0 dx x^{d/z-2}$ converges if

$$d/z - 2 > -1 \quad \Rightarrow \boxed{d > z}$$

BE condensation if $d > z$

c) critical temperature when $N = (N_{\max})$ (2)

$$N = V \mathcal{B} (k_B T)^{d/2} \cdot \mathcal{I}$$

$$k_B T_c = \left(\frac{N}{V \mathcal{B} \mathcal{I}} \right)^{2/d}$$

d) Below T_c : $\mu = 0$, condensate does not contribute to $\langle E \rangle, \langle p \rangle$

$$\begin{aligned} \langle E \rangle &= \int d\varepsilon g(\varepsilon) \frac{\varepsilon}{e^{\beta\varepsilon} - 1} = V \mathcal{B} (k_B T)^{d/2+1} \int_0^\infty dx \frac{x^{d/2}}{e^x - 1} \\ &= V \mathcal{B} (k_B T)^{d/2+1} \cdot \mathcal{I}_2 \end{aligned}$$

$$C_V \sim \frac{\partial \langle E \rangle}{\partial T} = \left(\frac{d}{2} + 1 \right) \frac{\langle E \rangle}{T} \sim T^{d/2}$$

$$e) \quad S = \int_0^T dT' \frac{C_V}{T'} = \frac{2}{d} \left(\frac{d}{2} + 1 \right) \frac{\langle E \rangle}{T} = \left(1 + \frac{2}{d} \right) \frac{\langle E \rangle}{T}$$

$$A = \langle E \rangle - TS = -\frac{2}{d} \langle E \rangle$$

$$P = \left(-\frac{\partial A}{\partial V} \right)_T = \frac{2}{d} \langle E \rangle / V \sim T^{d/2+1}$$

Problem 10.2

$$u(\varepsilon) = \frac{1}{\pi^2 h^3 c^3} \frac{\varepsilon^3}{e^{\beta\varepsilon} - 1}$$

maximum

$$\frac{d}{d\varepsilon} \frac{\varepsilon^3}{e^{\beta\varepsilon} - 1} = \frac{3\varepsilon^2}{e^{\beta\varepsilon} - 1} - \frac{\beta\varepsilon^3 e^{\beta\varepsilon}}{(e^{\beta\varepsilon} - 1)^2}$$

$$= \frac{3\varepsilon^2(e^{\beta\varepsilon} - 1) - \beta\varepsilon^3 e^{\beta\varepsilon}}{(e^{\beta\varepsilon} - 1)^2} = 0$$

$$3e^{\beta\varepsilon} - 3 = \beta\varepsilon e^{\beta\varepsilon}$$

Solve $3(e^x - 1) = x e^x \Rightarrow x = 2.8214$

$$\varepsilon = 2.8214 k_B T = 1.17 \cdot 10^{-22} \text{ Ws}$$

$$\varepsilon = h\nu = 2\pi h\nu = \frac{2\pi h c}{\lambda} \quad \omega = 1.1 \cdot 10^{12} \text{ s}^{-1}$$

$$\lambda = \frac{2\pi h c}{\varepsilon} = 1.7 \text{ mm}$$

Problem 10.3

Radiation of Betelgeuse

a) energy density $u(\epsilon)$

$$u(\epsilon) = \frac{1}{\pi^2 c^3 h^3} \frac{\epsilon^3}{e^{\beta\epsilon} - 1}$$

maximum at $\frac{du}{d\epsilon} = 0$

$$\frac{d}{d\epsilon} \frac{\epsilon^3}{e^{\beta\epsilon} - 1} = 0 \quad \Rightarrow \quad \frac{d}{dx} \frac{x^3}{e^x - 1} = 0$$

$$\frac{3x^2}{e^x - 1} - \frac{x^3 e^x}{(e^x - 1)^2} = 0 \quad \Rightarrow \quad 3x^2(e^x - 1) = x^3 e^x$$

$$(3-x)e^x = 3 \quad \text{numerical solution } x \approx 2.82$$

$$\beta\epsilon_{\text{max}} = 2.82 \quad \frac{1}{T_{\text{Bet}}} = \frac{\epsilon_{\text{max}}}{2.82 h\nu} \approx 3290 \text{ K}$$

b) $P_{\text{Bet}} \approx 10^4 P_{\text{sun}} = 10^4 \times 3.85 \times 10^{26} \text{ W} = 3.85 \times 10^{30} \text{ W}$

Stefan-Boltzmann:

$$P_{\text{Bet}} = 4\pi R_{\text{Bet}}^2 \times \frac{\pi^2}{60} \frac{(k_B T)^4}{h^3 c^2}$$

$$R_{\text{Bet}} = \sqrt{\frac{15}{\pi^3} \frac{h^3 c^2}{(k_B T)^4} P_{\text{Bet}}} = 2.15 \times 10^{11} \text{ m}$$

c) maximum of spectrum is in the red color range

radius much larger than sun