

9.1

a)

S = 0 bosons

$$n_a = 2 \quad n_b = 0 \quad \bar{E} = 2\varepsilon$$

$$n_a = 1 \quad n_b = 1 \quad \bar{E} = 0$$

$$n_a = 0 \quad n_b = 2 \quad \bar{E} = -2\varepsilon$$

partition function

$$Q = e^{-\beta 2\varepsilon} + 1 + e^{\beta 2\varepsilon}$$

$$Q = 1 + 2 \cosh(2\beta\varepsilon)$$

$$A = -k_B T \ln Q = -k_B T \ln(1 + 2 \cosh(2\beta\varepsilon))$$

$$\langle \bar{E} \rangle = -\frac{\partial \ln Q}{\partial \beta} = \frac{-4\varepsilon \sinh(2\beta\varepsilon)}{1 + 2 \cosh(2\beta\varepsilon)}$$

$$S = \frac{1}{T} (\bar{E} - A) = k_B \left[-\frac{4\beta\varepsilon \sinh(2\beta\varepsilon)}{1 + 2 \cosh(2\beta\varepsilon)} + \ln(1 + 2 \cosh(2\beta\varepsilon)) \right]$$

S = 1/2 fermions only $n_a = 1, n_b = 1$ statePauli: only $n_a = 1, n_b = 1$ state
allowed

$$Q = 1 \quad A = 0$$

$$\bar{E} = 0 \quad S = 0$$

b) $S=0$ bosons

$$n_a=2 \quad n_b=0 \quad \bar{E}=2\varepsilon$$

$$n_a=1 \quad n_b=1 \quad \bar{E}=\mu$$

$$n_a=0 \quad n_b=2 \quad \bar{E}=-2\varepsilon$$

Canonical probabilities

$$P_1 = \frac{e^{-\beta 2\varepsilon}}{e^{\beta 2\varepsilon} + e^{-\beta 2\varepsilon} + e^{-\beta \mu}}$$

$$P_2 = \frac{e^{-\beta \mu}}{e^{\beta 2\varepsilon} + e^{-\beta 2\varepsilon} + e^{-\beta \mu}}$$

$$P_3 = \frac{e^{\beta 2\varepsilon}}{e^{\beta 2\varepsilon} + e^{-\beta 2\varepsilon} + e^{-\beta \mu}}$$

positive μ suppresses state $(n_a=1, n_b=1)$
negative μ prefers state $(n_a=1, n_b=1)$

- for $\mu \rightarrow -\infty$ this state becomes ground state

- for $\mu \rightarrow \infty$ state goes to $\bar{E} \rightarrow \infty$ and drops out of thermal dynamics

9.2 Quantum corrections to classical ideal gas

The classical (Boltzmann) limit corresponds to

$$\langle n \rangle = \frac{1}{e^{\beta(\epsilon-\mu)} + \delta} \ll 1$$

Expanding the denominator gives

$$\langle n \rangle = e^{-\beta(\epsilon-\mu)}(1 - \delta e^{-\beta(\epsilon-\mu)}) + O[(e^{-\beta(\epsilon-\mu)})^3]$$

Average particle number and internal energy read

$$\begin{aligned} N &= \int d\epsilon a(\epsilon) e^{-\beta(\epsilon-\mu)}(1 - \delta e^{-\beta(\epsilon-\mu)}) \\ U &= \int d\epsilon \epsilon a(\epsilon) e^{-\beta(\epsilon-\mu)}(1 - \delta e^{-\beta(\epsilon-\mu)}) \end{aligned}$$

with

$$a(\epsilon) = \frac{V}{2\pi^2} \left(\frac{m}{\hbar^2} \right)^{3/2} \sqrt{2\epsilon}$$

The energy per particle is

$$\begin{aligned} \frac{U}{N} &= \frac{\int d\epsilon \epsilon^{3/2} e^{-\beta\epsilon}(1 - \delta e^{-\beta(\epsilon-\mu)})}{\int d\epsilon \epsilon^{1/2} e^{-\beta\epsilon}(1 - \delta e^{-\beta(\epsilon-\mu)})} \\ &= k_B T \frac{(3/4)\sqrt{\pi} (1 - \delta e^{\beta\mu}/2^{5/2})}{(1/2)\sqrt{\pi} (1 - \delta e^{\beta\mu}/2^{3/2})} \\ &= \frac{3}{2} k_B T (1 + \delta e^{\beta\mu} \sqrt{2}/8) \end{aligned}$$

The factor $e^{\beta\mu}$ in the correction term can be calculated in lowest order, i.e. for the classical ideal gas.

$$e^{\beta\mu} = \frac{N}{V} \left(\frac{2\pi\hbar^2}{mk_B T} \right)^{3/2} = \frac{N}{V} \lambda^3$$

where λ is the thermal wave length. Therefore

$$\frac{U}{N} = \frac{3}{2} k_B T (1 + \delta N \lambda^3 \sqrt{2}/(8V))$$